## Gauge-Invariant Localization of Infinitely Many Gravitational Energies from All Possible Auxiliary Structures

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#### Energy Localization Debate from 1910s

- Pseudotensor? 1916 (Einstein, 1923) for vs. (Schrödinger, 1918; Bauer, 1918) against.
- Reviews (Fletcher, 1960; Trautman, 1962; Goldberg, 1980; Szabados, 2004).
- ►  $\mathfrak{T}^{\mu}_{\nu}\sqrt{-g} =_{def} T^{\mu}_{\nu}\sqrt{-g} + t^{\mu}_{\nu}\sqrt{-g}$  satisfies  $\frac{\partial}{\partial x^{\mu}}(\mathfrak{T}^{\mu}_{\nu}\sqrt{-g}) = 0$  if (and only if)  $G^{\mu}_{\nu} = T^{\mu}_{\nu}$ . Here  $t^{\mu}_{\nu}\sqrt{-g}$  is gravity pseudotensor.
- Problem is not lack of conserved quantities, but too many and lack of relationships among them (Anderson, 1967).
- ► Pseudotensors aren't tensors, not even geometric objects.
- ► Many (1st derivatives) can vanish at any point or worldline.
- Mix of coordinate artifacts and real physics, it seems.

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#### References

#### Tensorial Energy from Additional Background Structure?

- ► Flat background metric (Rosen, 1963; Cornish, 1964).
- ► Coordinate change  $g_{\sigma\rho} \to e^{\pounds_{\xi}} g_{\sigma\rho}, u \to e^{\pounds_{\xi}} u, \eta_{\mu\nu} \to e^{\pounds_{\xi}} \eta_{\mu\nu}.$ *u* is matter (Grishchuk et al., 1984; Petrov, 2008).
- ► Problem of coordinate dependence reappears as gauge dependence: g<sub>σρ</sub> → e<sup>£</sup>ξ g<sub>σρ</sub>, u → e<sup>£</sup>ξ u, η<sub>µν</sub> → η<sub>µν</sub>.
- ▶  $\eta_{\mu\nu} \rightarrow e^{\pounds_{\xi}} \eta_{\mu\nu}$  affects energy density but no physical meaning.
- ► Flat connection  $\Gamma^{\alpha}_{\mu\nu} \rightarrow e^{\pounds} \epsilon \Gamma^{\alpha}_{\mu\nu}$  (Sorkin, 1991; Fatibene and Francaviglia, 2003) has analogous problem.
- ▶ Orthonormal tetrad  $e^{\mu}_{A}$ : extra local O(3,1) (Møller, 1964).
- Just moving lump in carpet, not flattening it out, with a flat metric, connection or tetrad: gauge-dependent energy.

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## Traditional Coping Mechanism

- Blame the question and invoke equivalence principle ad hoc: "[a]nybody who looks for a magic formula for 'local gravitational energy-momentum' is looking for the right answer to the wrong question." (Misner et al., 1973, p. 467)
- But Noether's theorems don't care about the equivalence principle. Maybe a different kind of invariance fits what Noether yields?
- Value of messy math as opposed to geometrical shortcuts and picture-thinking (Kiriushcheva et al., 2008; Brown, 2005).
- $\blacktriangleright$   $\infty$  many conserved energies (Bergmann, 1958; Komar, 1959).
- ► Drop *unique* localization, find new kind of co/invariance?

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#### No Need to Assume Energy is Unique

- Universal tacit undefended assumption of just one energy: 10 (or 16) components in one chart should suffice locally.
- ► (Goldberg, 1980) (Szabados, 2004, section 3.1.3); (Faddeev, 1982): "The energy of the gravitational field is not localized, i.e., a uniquely defined energy density does not exist."
- ▶ But no reason to believe in uniqueness—just a habit, default.
- ▶ ∞ many energies (Bergmann, 1958; Komar, 1959; Anderson, 1967; Regge and Teitelboim, 1974; Fatibene and Francaviglia, 2003). Any vector field gives one! Why can't they all be real?
- Uniqueness should have been rejected, localization could have been discovered 50 years ago by Bergmann.

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## Gauge Independence of Energies from All Background Structures: Case of Flat Metrics

- ► Don't choose one background; choose them all!  $\{(\forall \eta_{\rho\sigma}) \eta_{\rho\sigma}\}$ .
- Gauge invariance: nothing depends on arbitrary choice of  $\eta_{\rho\sigma}$ .
- ► Infinite-component invariant energy { $(\forall \eta_{\rho\sigma}) t^{\mu\nu}[g_{\alpha\beta}, \eta_{\rho\sigma}]$ }, each conserved:  $\partial_{1\mu}(\sqrt{-g}T^{\mu\nu} + \sqrt{-g}t^{\mu\nu}[g, \eta_1]) = 0$ ,  $\partial_{1\mu}\eta_{1\alpha\beta} \equiv 0$ ,  $\partial_{2\mu}(\sqrt{-g}T^{\mu\nu} + \sqrt{-g}t^{\mu\nu}[g, \eta_2]) = 0$ ,....
- ► Flat connection is analogous, but angular momentum problems (Chang et al., 2000) (c.f.) (Goldberg, 1958).
- {(∀η<sub>ρσ</sub>) η<sub>ρσ</sub>} a bit like group averaging (Marolf, 2002), but only collect into set, not add, gauge-dependent elements.

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#### Critique of Komar's Vector, Møller's Orthonormal Tetrad

- ► Komar's conserved quantities tensorial, depend on vector field.
- Problem: factor of 2 wrong answers for Komar (Katz, 1985; Katz and Ori, 1990; Iyer and Wald, 1994; Petrov and Katz, 2002; Fatibene and Francaviglia, 2003).
- ► Møller's orthonormal tetrad: accept them all?
- ► Worry: local Lorentz O(3, 1) group is gratuitous, so gauge invariance from *all tetrads* is Pickwickian.
- ► Flat metric  $\eta_{\rho\sigma}$ , flat connection  $\Gamma^{\alpha}_{\rho\sigma}$  give invariant localization.
- ►  $\eta_{\rho\sigma}$  or  $\Gamma^{\alpha}_{\rho\sigma}$  needed in action *S* (Faddeev, 1982; Hawking and Horowitz, 1996; Fatibene and Francaviglia, 2003).
- Gauge-invariant multi-action  $S[g_{\mu\nu}, \{(\forall \eta_{\rho\sigma}) \ \eta_{\rho\sigma}\}].$

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#### References

## Spinorial Almost Geometric Objects, So No Exception

- ► Spinors in coordinates (Ogievetskii and Polubarinov, 1965; Bilyalov, 2002; Gates et al., 1983):  $\langle r_{\mu\nu}, \psi \rangle$ , where  $r_{[\mu\alpha]} = 0$ ,  $g_{\mu\nu} = \sum_{\alpha=A} \sum_{\beta=B} r_{\mu\alpha} \eta^{AB} r_{\beta\nu}$ : symmetric square root.
- $\langle r_{\mu\nu}, \psi \rangle$  nonlinear spinor representation of coordinate transformations (up to sign), linear for conformal subgroup.
- ► No tetrad, no local Lorentz group;  $r_{\mu\alpha}$  locally like tetrad in symmetric gauge.
- ► Any coordinates, but swap to get time first (Bilyalov, 1992).
- ►  $\langle g_{\mu\nu}, \psi \rangle$  equivalent to  $\langle r_{\mu\nu}, \psi \rangle$ . Lie, covariant derivatives defined (Ogievetskii and Polubarinov, 1965; Szybiak, 1966).
- ► So energy localization proposed here suits spinors also.

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#### Pseudotensor in All Coordinates as Invariant Localization

- ► Tensors  $t^{\mu\nu}$  from  $\eta_{\rho\sigma}$  or  $\Gamma^{\alpha}_{\rho\sigma}$ : clear transformation properties.
- Ignoring global issues, can fix coordinates: each flat metric in Cartesian coordinates: η<sub>RS</sub> = diag(-1, 1, 1, 1), Γ<sup>A</sup><sub>RS</sub> = 0.
- ▶ Yields pseudotensor in  $g_{\mu\nu}$  in every gauge/coordinate system.
- Invariant: for every chart U,  $\{\forall U t^{\mu}_{\nu}[g_{\mu\nu}, \eta_{MN}]\}$ .
- ► A *pseudotensor* is *okay*; just take it in *all* coordinate systems.
- ► Not ∞ many faces of same entity as with tensor, but ∞ many distinct entities, each in own adapted coordinate system.
- ► Not just metric, but natural bases enter definition of energy.
- ► Good pseudotensor maybe same for all solutions (Katz et al., 1997), maybe not (Nester, 2004).

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#### Objections to Pseudotensors Assume Unique Energy

- Worry:  $t^{\mu}_{\nu}$  vanishing at point or worldline in some coordinates.
- ▶ Reply: some but not all energies vanish there.
- Worry: Minkowski in unimodular spherical coordinates has nonzero Einstein pseudotensor (Bauer, 1918; Pauli, 1921).
- ► Reply: different coordinate system gives different energies.
- ► Worry: total energy in these coordinates diverges.
- ► Reply: spherical coordinate singularities, strongly curved basis.
- ▶ Worry: Einstein pseudotensor 0 [for r > 2M] in Schwarzschild in  $\approx$  Cartesian  $\sqrt{-g} = 1$  coordinates (Schrödinger, 1918).
- Reply: Many energies exist; some vanish outside horizon, some don't (Petrov, 2005; Petrov, 2008).

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#### References

#### Noether Operator Is Invariant in New Sense

- Non-GR fields: ∇<sub>μ</sub>(T<sup>μν</sup>√-gξ<sub>ν</sub>) = (T<sup>μν</sup>√-gξ<sub>ν</sub>),<sub>μ</sub> = 0: conserved vector density T<sup>μν</sup>√-gξ<sub>ν</sub> is algebraic in ξ<sup>ν</sup>.
- For GR, Noether's theorem gives nontensorial differential operator in ξ<sup>ν</sup> (Schutz and Sorkin, 1977; Sorkin, 1977; Thirring and Wallner, 1978; Szabados, 1992).
- ► Feeding it natural basis yields pseudotensor components.
- Feeding it all natural bases yields pseudotensor in all coordinates: invariant in sense proposed here.
- Not unique due to possibility of adding curls.
- But even scalar fields in flat space-time have a bit of that problem (Callan et al., 1970): "improved" stress tensor.

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# Logical Equivalence of All Conservation Laws to Einstein's Equations

- Fields: conservation because every field has Euler-Lagrange equations or (generalized) Killing vectors (Trautman, 1966).
- GR (without  $\eta_{\rho\sigma}$  or  $\Gamma^{\alpha}_{\rho\sigma}$ ): every field has E-L equations.
- ► GR: conservation from Einstein's equations without using matter equations (Anderson, 1967; Wald, 1984).
- ► Coordinate form  $\frac{\partial}{\partial x^{\mu}}(\mathfrak{T}^{\mu}_{\nu}\sqrt{-g})=0$  in all coordinates.
- Sheds light on relation of GR to first law of thermodynamics: GR obviously entails it.

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# Logical Equivalence of All Conservation Laws to Einstein's Equations

- Reverse entailment also holds (Anderson, 1967): pseudotensor law in all coordinates entails Einstein's equations!
- Illuminates spin-2 derivations of Einstein's equations (Einstein and Grossmann, 1996; Deser, 1970; Pitts and Schieve, 2001).
- Clearly, nothing logically equivalent to Einstein's equations depends viciously on coordinates.
- ► Thus pseudotensor laws don't depend viciously on coordinates.
- ► Pseudotensor laws give invariant localization of energy.
- ► Another way to see that pseudotensor localization is real.

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## Angular Momentum Localization

- ► Generalization straightforward: invariant ∞-component angular momentum. Depends on x<sup>µ</sup> explicitly.
- ► Matrix diag(-1,1,1,1) helps with angular momentum (Chang et al., 2000) (c.f.) (Goldberg, 1958).
- ► Symmetric total energy-momentum  $\sqrt{-g}\mathfrak{T}^{\mu\nu}$  gives angular momentum complex  $\mathfrak{M}^{\mu\nu\alpha} =_{def} \sqrt{-g}\mathfrak{T}^{\mu\nu}x^{\alpha} \sqrt{-g}\mathfrak{T}^{\mu\alpha}x^{\nu}$ .
- $\blacktriangleright \ \frac{\partial}{\partial x^{\mu}}\mathfrak{M}^{\mu\nu\alpha} = 0 \text{ due to } \frac{\partial}{\partial x^{\mu}}(\sqrt{-g}\mathfrak{T}^{\mu\nu}) = 0 \text{ and } \mathfrak{T}^{[\mu\nu]} = 0.$
- Choice of pseudotensor (Katz et al., 1997; Nester, 2004) affects distributions of angular momenta.
- Using  $\{(\forall \eta_{\rho\sigma}) \ \eta_{\rho\sigma}\}$ , position 4-vectors replace  $x^{\mu}$ .

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#### Conceptual Benefits of Local Energy Conservation

Lack of local energy conservation in GR in general or Big Bang cosmology has been invoked for unwarranted conclusions, such as:

- GR is false (Logunov and Folomeshkin, 1977; Logunov et al., 1986)—addressed in (Faddeev, 1982; Zel'dovich and Grishchuk, 1988; Grishchuk, 1990);
- Big Bang cosmology is false (by Robert Gentry)—addressed in (Pitts, 2004a; Pitts, 2004b);
- Big Bang cosmology is plausibly true and yet violates a principle so fundamental as to transcend physics into metaphysics (Bunge, 2000);

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- Big Bang cosmology is a heat sink for anomalous terrestrial heat (by D. Russell Humphreys)—addressed in (Pitts, 2009b);
- GR makes it easier than other theories for souls to affect bodies (Collins, 2008)—addressed in (Pitts, 2009a);
- 6. Universes with zero total energy can pop into being without violating energy conservation (Tryon, 1973; Thirring, 2003).
- Concerning 6, all energy densities needed for gauge invariance must vanish; impossible except maybe in boring cases.
- Detailed analysis undermines these 6 conclusions, if one tries.
  But new claims keep arising.
- Play with matches and keep fire extinguishers handy?
- Such claims would be unthinkable in ordinary field theories due to well-known gauge-invariant local conservation laws.

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#### Conclusions

- Ironic that Einstein took energy conservation as criterion for GR equations, yet GR is widely held to have no such law.
- $\blacktriangleright$  Irony resolved by  $\infty\text{-component}$  gauge-invariant localization.
- ► Natural bases, not just metric, determine energies.
- ► GR logically equivalent to conservation of ∞ many energies, hence *more* conserving of energy than other theories.
- Gauge (in)dependence largely orthogonal to question of 'right answers' for the conserved quantities.
- Non-uniqueness of relocalizing (Anderson, 1967), as in other field theories. Maybe case-by-case (Nester, 2004).
- ▶ Best functional form technically, make covariant in above way.

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