

Introductory Astrophysics Teaching Manual

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August 7, 2024

1 Observational Astrophysics

1. Engage

- Learning objectives: Today we are going to learn about how scientists discover and study planets outside our solar system.
- Key words: exoplanet, flux, period, transit depth, folding
- Knowledge Outcomes:
 - How scientists observe exoplanets
 - What exoplanet data looks like
 - How scientists use plots to
 - How exoplanet size can be estimated from flux data
- Skill Outcomes:
 - Use a jupyter notebook to run python code
 - Make value estimates by looking at plots

2. Prior Learning

- Check knowledge about our own solar system
 - Number of planets
 - Different planet sizes and compositions
 - Planets orbit the Sun
- Discussion of telescopes and space observatories as background for “usual” way of observing space objects

3. Intro New Knowledge

- Define **exoplanets** and discuss categories of them in relation to the planets in our solar system
- Differentiate what are “usual” observation methods with how exoplanets typically observed - **transit method** (call back to introduction)
- *Physical Demo 1*: transits - hold a ball and spin in a circle?

4. Build New Content

- **Question:** Given the transit method demo, what do you expect a graph of the total light from a system to look like? (Keep using physical demo to get students engaged)
- Details of transit method - **period, transit depth**
- Emphasize what the signal looks like (graphic and call back to physical demo 1)

5. Apply Learning

- “Students have the opportunity to apply what they have learnt to new situations.”
- *Investigation:* Students will explore and attempt to answer a series of questions.
 - (a) How can we estimate the period of a transit?
 - (b) How can we estimate the transit depth of the exoplanet?
 - (c) Physical demo 2: folding - baking paper?
 - (d) What happens to the folded light curve when you use your estimate of the period versus when you use the known period of the exoplanet?
 - (e) What type of exoplanet would you classify this as?

2 Theoretical Astrophysics

1. Engage

- Learning objectives: Today we are going to learn about how scientists discover and study mysterious objects called black holes.
- Key words: binary black hole, gravitational waves, strain, amplitude
- Knowledge Outcomes:
 - Understand how scientists observe black holes (it is different than observing EM sources)
 - Signals carry information about their sources
 - Fitting theoretical signal templates to real data gives scientific insight
- Skill Outcomes:
 - Use a jupyter notebook to run python code
 - Change input values to match fit with data

2. Prior Learning

- Check knowledge about stellar life cycle - what happens to stars when they die?
- **Question:** What do you know about black holes?

3. Intro New Knowledge

- Define **black holes** and how they can be in binary systems
- Discuss how space is like a fabric called “spacetime” - really heavy objects distort spacetime in a measurable way, called **gravitational waves**
- *Video:* LIGO detecting gravitational waves
- LIGO is measuring a quantity called **strain**

4. Build New Content

- **Question:** What do you think we can learn from the signals?
- Shapes of signals tell us information about the sources
- Discuss theoretical signal waveforms and real data
- Mass and distance and their effects on **amplitude** - emphasize this relationship

5. Apply Learning

- “Students have the opportunity to apply what they have learnt to new situations.”
- *Investigation:* Students will explore and attempt to answer a series of questions.
 - (a) What does the signal look like when the black holes have similar or the same mass?
 - (b) What happens if you make the masses really large?
 - (c) What happens if you make the masses really different from each other?
 - (d) What happens if you make the distance large (i.e. far away)?

3 Astronomical Instrumentation

1. Engage

- Learning objectives: Today we are going to learn about how scientists take and analyze images to study objects in space.
- Key words: photovoltaic effect, image effects, stars, galaxies
- Knowledge Outcomes:
 - General idea of how detectors convert observed photons into a digital signal
 - Understanding that image cleaning is required to produce images shown in media
- Skill Outcomes:
 - Identifying morphological differences between stars and galaxies in a real JWST image
 - Differentiating between observations vs. artifacts in images

2. Prior Learning

- Check misconceptions about astronomy images displayed in media. Do students think images come out perfectly when taken? Allude to image cleaning without explicitly mentioning it.
- Ask the students if they know how digital cameras create images

3. Intro New Knowledge

- Recap: everything in the Universe is made of atoms, and atoms contain electrons
- Some special materials can help us use these electrons to convert light into a digital signal
- *Whiteboard demo*: draw a star inscribed into a 3x3 pixel grid, then a 10x10 pixel grid to introduce concept of resolution. Relate to Figure 12 in lab manual

4. Build New Content

- **Question**: How do we expect the brightness of something to correlate to how many electrons it knocks loose in the detector?
- More electrons = brighter pixels, more pixels = better resolution, therefore we can get a really nice image with a larger detector!

5. Apply Learning

- “Students have the opportunity to apply what they have learnt to new situations.”
- *Investigation*: Students will explore and attempt to answer a series of questions.
 - (a) How many of your objects are galaxies and how many are stars? How can you tell?
 - (b) What could help you better differentiate between the two types of objects?
 - (c) Before you apply the image effects, take a second to think about what each of them might do to the image.
 - (d) After having gone through each effect, what can you say about what they do to the image? How many of your initial guesses were right?
 - (e) **CLOSEOUT CHALLENGE QUESTION**: How would you remove some of these effects from your images? Can they all be removed? Cost considerations, time considerations, etc.